

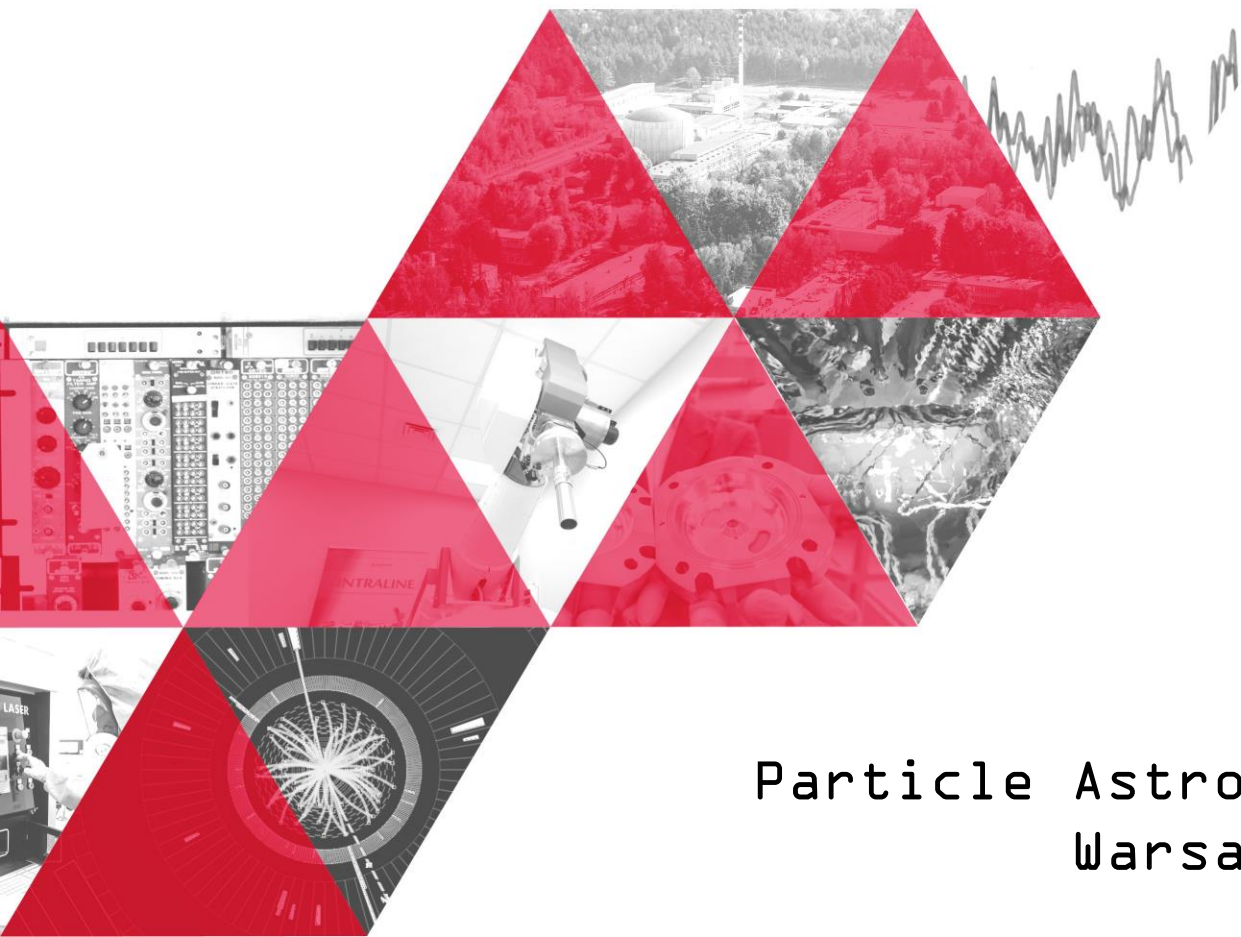
POLAR and POLAR-2 Instruments

Dominik Rybka

on behalf of the POLAR Collaboration



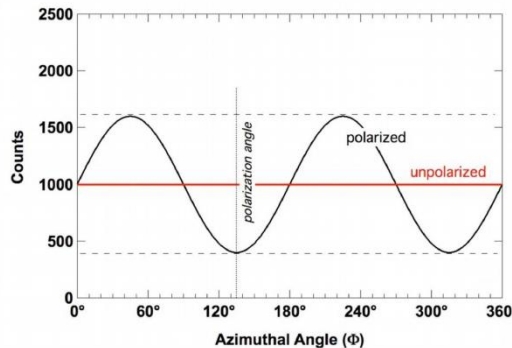
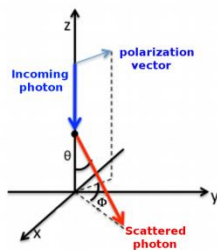
NARODOWE
CENTRUM
BADAŃ
JĄDROWYCH
ŚWIERK



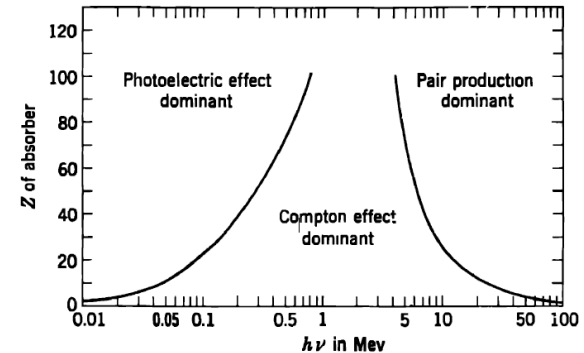
Particle Astrophysics in Poland
Warsaw, 20-21 May 2019

Gamma-Ray Burst Polarimetry

- Most energetic events in the universe since the big bang
- Timing, Direction and Energy spectrum measured in great detail
- Two parameters remain: polarisation degree and polarisation angle
- Polarisation holds information on emission process, e.g. synchrotron emission or photospheric emission, also on emission region.



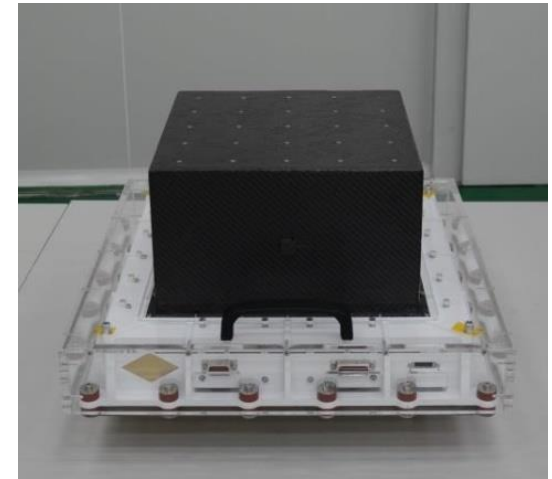
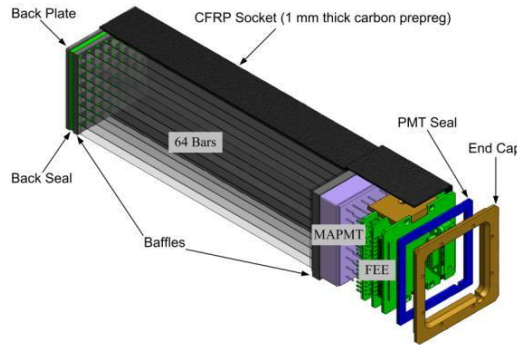
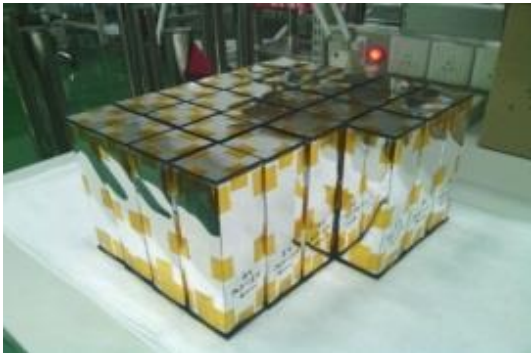
Azimuthal scattering angle dependence on polarisation



Three major types of photon interaction depending on the energy of the incoming photon and on the atomic number of the absorber

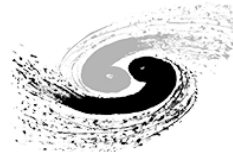
POLAR design

- segmented scintillator array to measure the Compton scattering angle
- in total 1600 plastic scintillators, $6 \times 6 \times 176 \text{ mm}^3$, EJ-248M, optimised for the cross section for Compton scattering in the 50-500 keV energy range
- plastic scintillators allow for a relatively large effective area, with low mass of 30kg
- each group of 64 scintillators is read-out using a single MAPMT H8500 from Hamamatsu
- relatively large effective area ($30 \times 30 \text{ cm}$)
- small pixels allows for high precision scattering angle measurements
- uniform effective area gives us a large Field of View (1/3rd of the sky)



POLAR Polish tasks

- Back End electronics design
- VHDL Programming of FEE and BEE
- High Voltage Power Supply design



Institute of High Energy Physics
Chinese Academy of Sciences

POLAR launch



Jiuquan Satellite Launch Centre
15 September 2016

POLAR results

LETTERS

<https://doi.org/10.1038/s41550-018-0664-0>

nature
astronomy

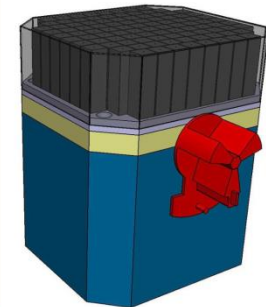
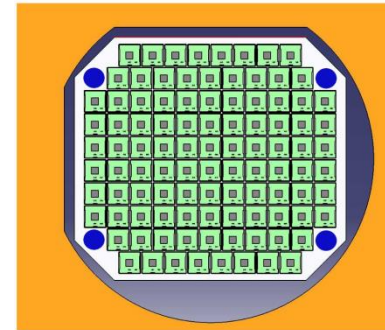
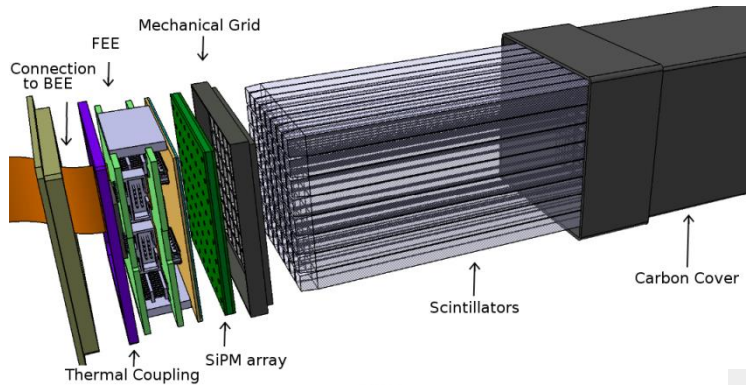
Detailed polarization measurements of the prompt emission of five gamma-ray bursts

Shuang-Nan Zhang ^{1,2,10*}, Merlin Kole ^{3,10*}, Tian-Wei Bao¹, Tadeusz Batsch⁴, Tancredi Bernasconi⁵, Franck Cadoux³, Jun-Ying Chai^{1,2}, Zi-Gao Dai^{6,7}, Yong-Wei Dong¹, Neal Gauvin⁵, Wojtek Hajdas⁸, Mi-Xiang Lan^{6,9}, Han-Cheng Li^{1,2}, Lu Li¹, Zheng-Heng Li ^{1,2}, Jiang-Tao Liu¹, Xin Liu^{1,2}, Radoslaw Marcinkowski⁸, Nicolas Produit ⁵, Silvio Orsi³, Martin Pohl³, Dominik Rybka⁴, Hao-Li Shi¹, Li-Ming Song^{1,2}, Jian-Chao Sun¹, Jacek Szabelski⁴, Teresa Tymieniecka⁴, Rui-Jie Wang¹, Yuan-Hao Wang^{1,2}, Xing Wen^{1,2}, Bo-Bing Wu¹, Xin Wu³, Xue-Feng Wu⁹, Hua-Lin Xiao^{1,8}, Shao-Lin Xiong¹, Lai-Yu Zhang¹, Li Zhang¹, Xiao-Feng Zhang¹, Yong-Jie Zhang¹ and Anna Zwolinska⁴

Nature Astronomy Vol. 3 No 1 (2019)

POLAR-2 design

- launch in 2014, on-board of Chinese Space Station
- 4 times bigger: 100 modules with 64 plastic scintillator bars of $6 \times 6 \times 176 \text{ mm}^3$
- SiPM arrays instead of MAPMT
- 4 additional spectrometers based on the GECAM detectors
- Field of View: half of the sky
- $580 \times 650 \times 770 \text{ mm}^3$, 160kg, 300W
- gain in sensitivity of a factor 9 (bigger instrument, new detectors)



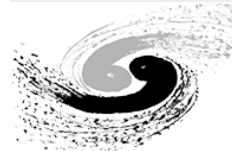
MAX PLANCK INSTITUTE
FOR EXTRATERRESTRIAL PHYSICS



Narodowe Centrum Badań Jądrowych
National Centre for Nuclear Research

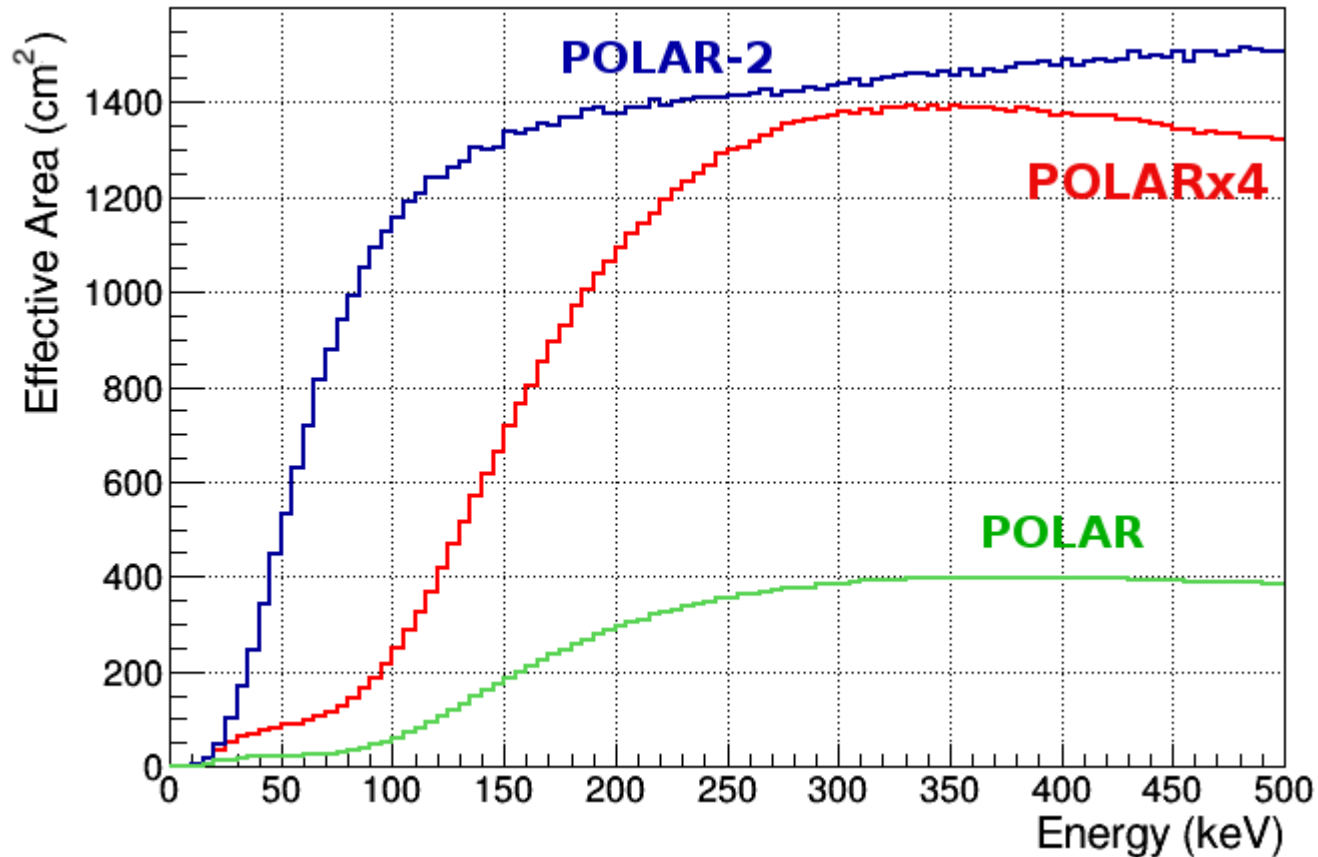
ŚWIERK

instytut kategorii A+, JRC collaboration partner



Institute of High Energy Physics
Chinese Academy of Sciences

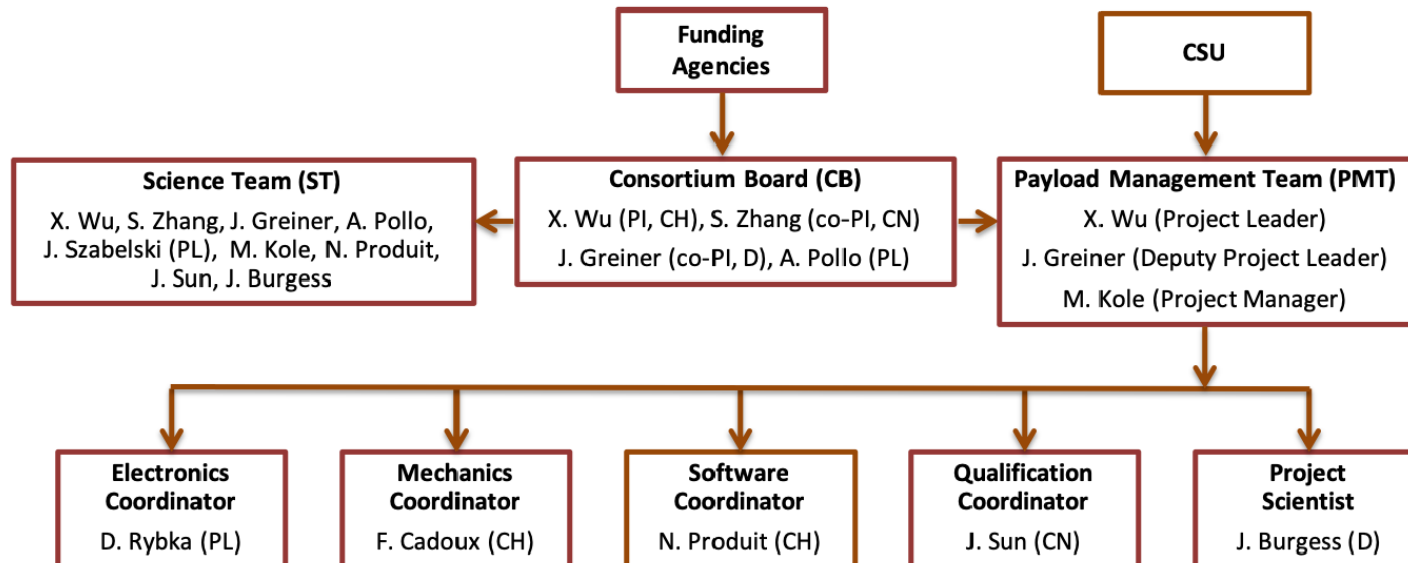
POLAR-2 gains at lower energies



Monte Carlo simulation of effective area of POLAR-2

POLAR-2 Polish tasks

- Front End electronics design and production
- Back End electronics design and production
- VHDL Programming of FEE and BEE
- Low Voltage Power Supply





Thank You

