

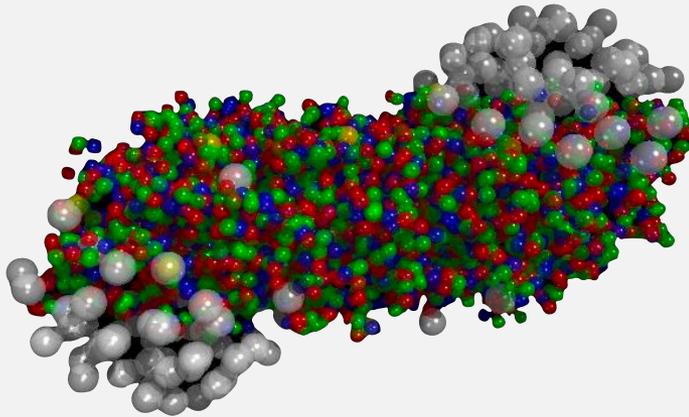
Quark-gluon plasma in magnetic fields

Patrycja Słoń

National Centre for Nuclear Research, Warsaw, Poland

Supervised by **Stanisław Mrówczyński**

The quark-gluon plasma (QGP)



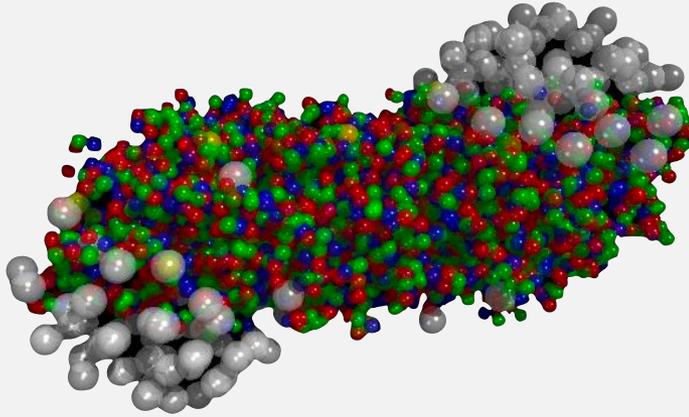
- Extremely hot (~ 300 MeV
 $\approx 3.5 \cdot 10^{12}$ K)

- Phase of matter consisting of deconfined quarks and gluons

mass	$\approx 2.16 \text{ MeV}/c^2$	$\approx 1.273 \text{ GeV}/c^2$	$\approx 172.57 \text{ GeV}/c^2$	0
charge	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0
spin	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1
	u	c	t	g
	up	charm	top	gluon
QUARKS	$\approx 4.7 \text{ MeV}/c^2$	$\approx 93.5 \text{ MeV}/c^2$	$\approx 4.183 \text{ GeV}/c^2$	
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	
	d	s	b	
	down	strange	bottom	

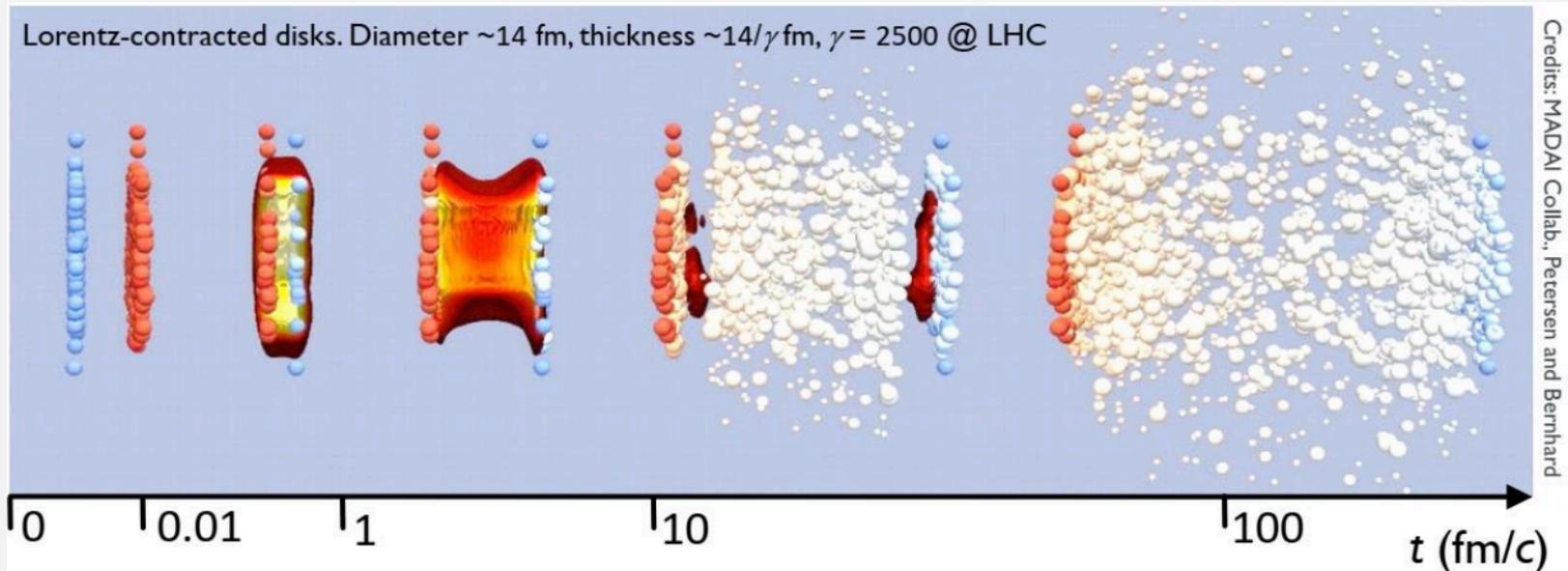
- Asymptotic freedom - the strong force between quarks and gluons decreases with increasing relative momentum
- Colour confinement - colour-charged particles cannot be isolated

The quark-gluon plasma (QGP)



- Extremely hot (~ 300 MeV $\approx 3.5 \cdot 10^{12}$ K)

- The QGP existed in nature only soon after the Big Bang (10^{-12} to 10^{-6} s)
- We are able to study the QGP in the laboratory through heavy-ion collision experiments

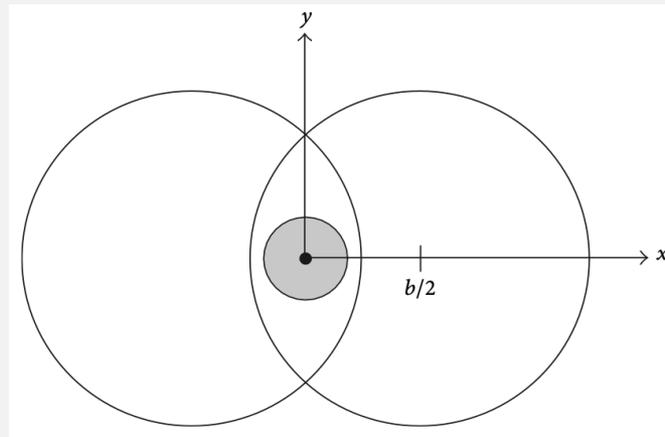


Units: t [fm/ c], $c = 1$, $\hbar = 1$

- Measuring time in fm/ c is convenient for describing the motion of high-speed particles in a laboratory setting.
- We understand it as time in which the light traveled a femtometer.
- In 1 second the light will travel $3 \cdot 10^8$ m
($1 \text{ s} \cdot c = 3 \cdot 10^8 \text{ m}$)
- The light will travel a 1 meter in $3.3 \cdot 10^{-9}$ s
($1 \text{ m}/c = 1/(3 \cdot 10^8) \approx 3.3 \cdot 10^{-9} \text{ s}$)
- The light will travel a 1 femtometer in $3.3 \cdot 10^{-24}$ s
($1 \text{ fm}/c = 10^{-15} \text{ m}/c = 10^{-15}/(3 \cdot 10^8) \approx 3.3 \cdot 10^{-24} \text{ s}$)
- A femtometer is a size characteristic for the problem of nuclei collision
- Nucleon size $\approx 10^{-15} \text{ m} = 1 \text{ fm}$

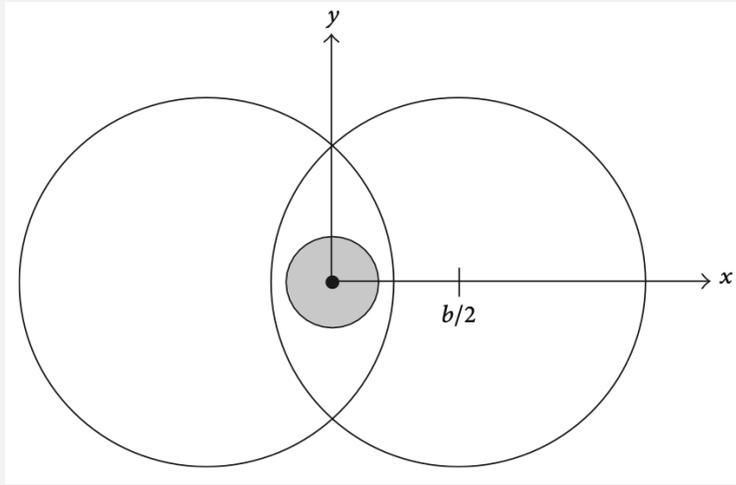
Magnetic field in the presence of strong interactions

- Quarks interact with each other via the strong interactions (100 times stronger than the electromagnetic force)
- We cannot neglect the magnetic field when it is of a great magnitude
- We estimate the magnetic field in heavy-ion collisions using the Biot-Savart law (equation describing the magnetic field generated by a constant electric current).



$$B \sim \gamma Z e \frac{b}{R^3}$$

Magnetic field in heavy-ion collisions



K. Tuchin, *Particle production in strong electromagnetic fields in relativistic heavy-ion collisions*,
Adv. High Energy Phys. 2013, 490495

Heavy-ion collision geometry as seen along the collision axis z

$$\text{Biot-Savart law: } B \sim \gamma Ze \frac{b}{R^3}$$

- R - the radius of two ions
- Ze - electric charge of the ions
- b - impact parameter
- $\gamma = \sqrt{s_{NN}}/2m_N$ - Lorentz factor.

$$eB \sim 10^{18} \text{ G} \approx m_\pi^2$$

- At RHIC heavy ions are collided at 200 GeV per nucleon, hence $\gamma = 100$. Using $Z = 79$ for gold and $b \sim R_A \approx 7$ fm.

Magnetic field in heavy-ion collisions

$$eB \sim 10^{18} \text{ G} \approx m_\pi^2 \approx 1.96 \cdot 10^4 \text{ MeV}^2$$

- The influence of electromagnetic field is especially noticeable when it is of a size characteristic for strong interactions. For the magnetic field it's when it's proportional to the square of pion mass.
- the strongest magnetic field created on earth in a form of electromagnetic shock wave: $\sim 10^7 \text{ G}$
- magnetic field of a neutron star: $10^{10} - 10^{13} \text{ G}$,
- magnetic field of a magnetar: up to 10^{15} G .

The magnetic field in a vacuum

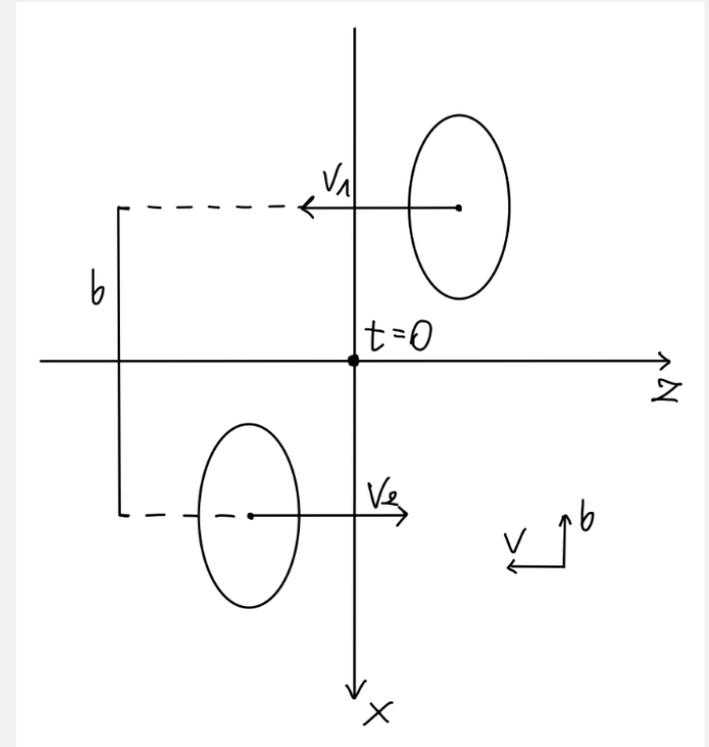
- the point of reference -

- Two identical nuclei collide at $t = 0$ with an impact parameter b

$$\begin{cases} \mathbf{r}_{01} = (b/2, 0, 0), \\ \mathbf{v}_1 = (0, 0, v), \\ \mathbf{R}_1 = (x - b/2, y, z - vt), \end{cases} \quad \begin{cases} \mathbf{r}_{02} = (-b/2, 0, 0), \\ \mathbf{v}_2 = (0, 0, -v), \\ \mathbf{R}_2 = (x + b/2, y, z + vt), \end{cases}$$

$$\mathbf{R} \equiv \mathbf{r} - \mathbf{r}_0 - \mathbf{v}t$$

- The problem was already studied and the magnetic field can be calculated using the Li'enard-Wiechert potentials



$$\mathbf{B}(t, \mathbf{r}) = \frac{q(1 - \mathbf{v}_1^2)\mathbf{v}_1 \times \mathbf{R}_1}{(\mathbf{R}_1^2 - (\mathbf{R}_1 \times \mathbf{v}_1)^2)^{3/2}} + \frac{q(1 - \mathbf{v}_2^2)\mathbf{v}_2 \times \mathbf{R}_2}{(\mathbf{R}_2^2 - (\mathbf{R}_2 \times \mathbf{v}_2)^2)^{3/2}}$$

The magnetic field in a vacuum - the point of reference -

$$\mathbf{B}(t, \mathbf{r}) = \frac{q(1 - v_1^2)\mathbf{v}_1 \times \mathbf{R}_1}{(\mathbf{R}_1^2 - (\mathbf{R}_1 \times \mathbf{v}_1)^2)^{3/2}} + \frac{q(1 - v_2^2)\mathbf{v}_2 \times \mathbf{R}_2}{(\mathbf{R}_2^2 - (\mathbf{R}_2 \times \mathbf{v}_2)^2)^{3/2}}$$

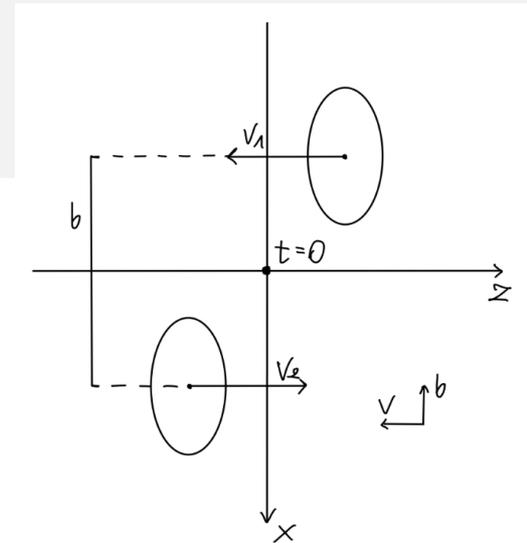
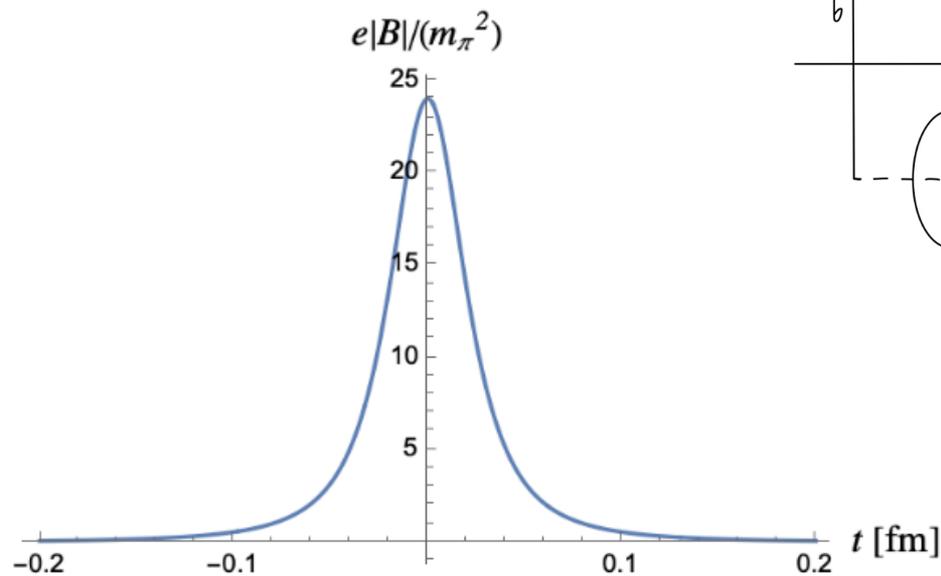


Fig. 2. Magnetic field as a function of time for Au-Au collision - reference

Electromagnetic field in heavy-ion collisions

- The Maxwell's equations

$$\left\{ \begin{array}{l} \nabla \cdot \mathbf{D}(t, \mathbf{r}) = 4\pi\rho_{ext}(t, \mathbf{r}), \\ \nabla \cdot \mathbf{B}(t, \mathbf{r}) = 0, \\ \nabla \times \mathbf{E}(t, \mathbf{r}) = -\frac{\partial \mathbf{B}(t, \mathbf{r})}{\partial t}, \\ \nabla \times \mathbf{B}(t, \mathbf{r}) = 4\pi\mathbf{j}_{ext}(t, \mathbf{r}) + \frac{\partial \mathbf{D}(t, \mathbf{r})}{\partial t}. \end{array} \right. \begin{array}{l} - \text{ Gauss' law} \\ - \text{ Gauss' law for magnetism} \\ - \text{ Faraday's law} \\ - \text{ Modified Ampère's law} \end{array}$$

- **D** - the electric displacement field

- $\frac{\partial \mathbf{D}(t, \mathbf{r})}{\partial t}$ - the displacement current - an additional source of the magnetic field. It is not an electric current of moving charges, but a time-varying electric field.

Dielectric permittivity tensor

- The electric displacement field \mathbf{D}

$$\mathbf{D}(\omega, \mathbf{k}) = \hat{\epsilon}(\omega, \mathbf{k})\mathbf{E}(\omega, \mathbf{k})$$

- $\hat{\epsilon}(\omega, \mathbf{k})$ - the dielectric permittivity tensor - provides information about the medium in which the heavy-nuclei collide
- The medium properties are described with (ω, \mathbf{k}) . It requires us to transform the Maxwell equations from (t, \mathbf{r}) to (ω, \mathbf{k}) dependence. We do that using the Fourier transformations.

Two-sided Fourier transformation

$$f(\omega, \mathbf{k}) \equiv \int_{-\infty}^{\infty} dt \int d^3r e^{i(\omega t - \mathbf{k} \cdot \mathbf{r})} f(t, \mathbf{r})$$

$$f(t, \mathbf{r}) = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \int \frac{d^3k}{(2\pi)^3} e^{-i(\omega t - \mathbf{k} \cdot \mathbf{r})} f(\omega, \mathbf{k})$$

- Assumes that the medium in which the collision happens doesn't undergo any changes from time $t = -\infty$ to $t = +\infty$.
- In reality the production of quark-gluon plasma at $t = 0$ modifies the electromagnetic field due to the quark currents.
- The use of the two-sided Fourier transformation implies that those currents exist even before the plasma is present.
- The two-sided Fourier transformation is not sufficient for the initial condition problem.

One-sided Fourier transformation

Magnetic field in reality

- Before the collision ($t < 0$) there is no plasma. Only the approaching nuclei generate the magnetic field.
- The additional fields generated due to the plasma appear at $t = 0$
- At $t > 0$ we see the effects of the quark-gluon plasma on the electromagnetic field.

$$f(\omega, \mathbf{k}) \equiv \int_0^{\infty} dt \int d^3r e^{i(\omega t - \mathbf{k} \cdot \mathbf{r})} f(t, \mathbf{r})$$

$$f(t, \mathbf{r}) = \int_{-\infty + i\sigma}^{\infty + i\sigma} \frac{d\omega}{2\pi} \int \frac{d^3k}{(2\pi)^3} e^{-i(\omega t - \mathbf{k} \cdot \mathbf{r})} f(\omega, \mathbf{k})$$

The real parameter $\sigma > 0$ is chosen in such a way that the integral over ω is taken along a straight line in the complex ω -plane, parallel to the real axis, above all singularities of $f(\omega, \mathbf{k})$.

The magnetic field in heavy-ion collisions

The two-sided Fourier transformation

$$f(\omega, \mathbf{k}) \equiv \int_{-\infty}^{\infty} dt \int d^3r e^{i(\omega t - \mathbf{k} \cdot \mathbf{r})} f(t, \mathbf{r})$$

$$f(t, \mathbf{r}) = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \int \frac{d^3k}{(2\pi)^3} e^{-i(\omega t - \mathbf{k} \cdot \mathbf{r})} f(\omega, \mathbf{k})$$

- The Maxwell's equations

$$\left\{ \begin{array}{l} \mathbf{k} \cdot \mathbf{D}(\omega, \mathbf{k}) = 4\pi\rho_{ext}(\omega, \mathbf{k}), \\ \mathbf{k} \cdot \mathbf{B}(\omega, \mathbf{k}) = 0, \\ \mathbf{k} \times \mathbf{E}(\omega, \mathbf{k}) = \omega\mathbf{B}(\omega, \mathbf{k}), \\ \mathbf{k} \times \mathbf{B}(\omega, \mathbf{k}) = -\omega\mathbf{D}(\omega, \mathbf{k}) - 4\pi i \mathbf{j}_{ext}(\omega, \mathbf{k}). \end{array} \right. \begin{array}{l} \text{- Gauss' law} \\ \text{- Gauss' law for magnetism} \\ \text{- Faraday's law} \\ \text{- Modified Ampère's law} \end{array}$$

- The displacement field \mathbf{D}

$$\mathbf{D}(\omega, \mathbf{k}) = \hat{\epsilon}(\omega, \mathbf{k})\mathbf{E}(\omega, \mathbf{k})$$

$$\mathbf{B}(\omega, \mathbf{k}) = -4\pi i \frac{\mathbf{k} \times \mathbf{j}_{ext}(\omega, \mathbf{k})}{\omega^2 \epsilon_T(\omega, \mathbf{k}) - \mathbf{k}^2}$$

- The current density \mathbf{j}_{ext}

$$\mathbf{j}_{ext}(t, \mathbf{r}) = q\mathbf{v}\delta(\mathbf{r} - \mathbf{r}_0 - \mathbf{v}t)$$

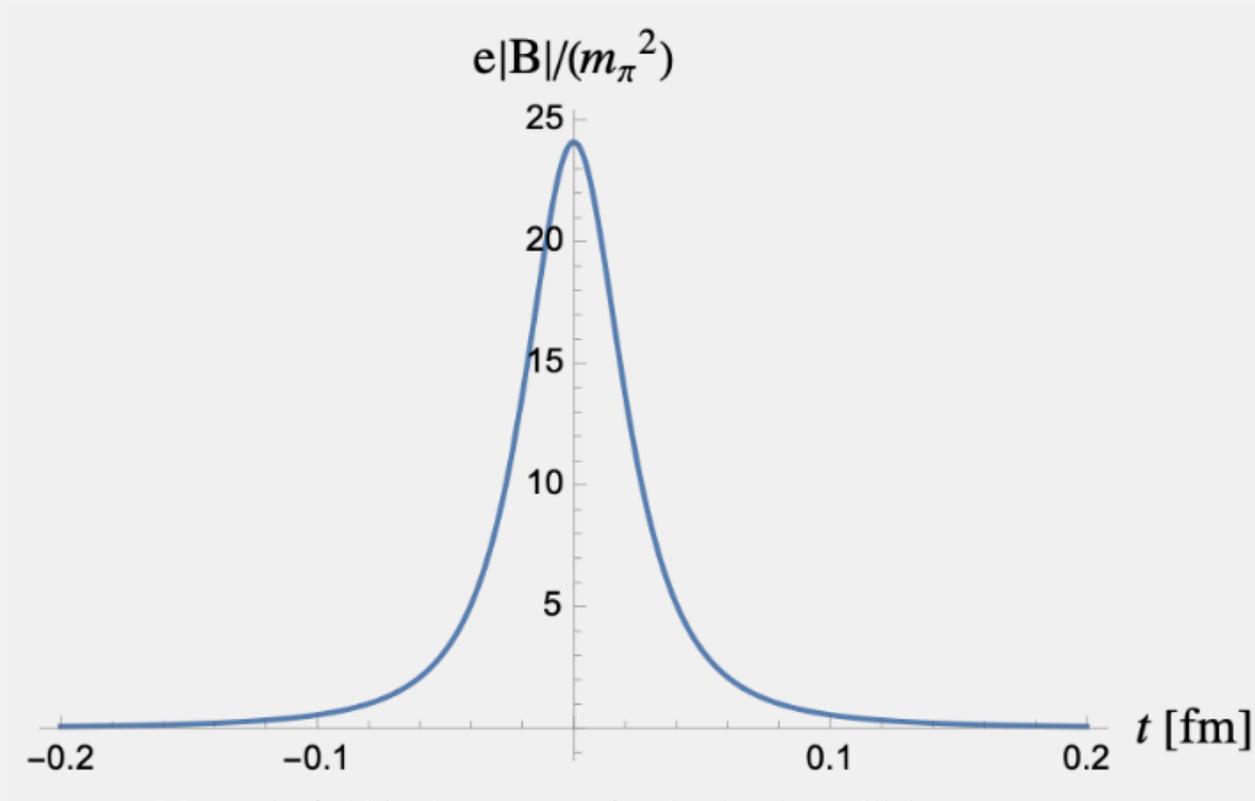
$$\mathbf{j}_{ext}(\omega, \mathbf{k}) = 2\pi q\mathbf{v} \delta(\omega - \mathbf{k} \cdot \mathbf{v}) e^{-i\mathbf{k} \cdot \mathbf{r}_0}$$

Magnetic field in the vacuum

$$\varepsilon_T = 1$$

$$\mathbf{B}(\omega, \mathbf{k}) = -4\pi i \frac{\mathbf{k} \times \mathbf{j}_{ext}(\omega, \mathbf{k})}{\omega^2 - \mathbf{k}^2}$$

$$\mathbf{B}(t, \mathbf{k}) = -4\pi i Z e \frac{\mathbf{k} \times \mathbf{v}}{(\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2} e^{-i(\mathbf{k} \cdot \mathbf{v}t + \mathbf{k} \cdot \mathbf{r}_0)}$$



Magnetic field in the vacuum for the Au-Au collision

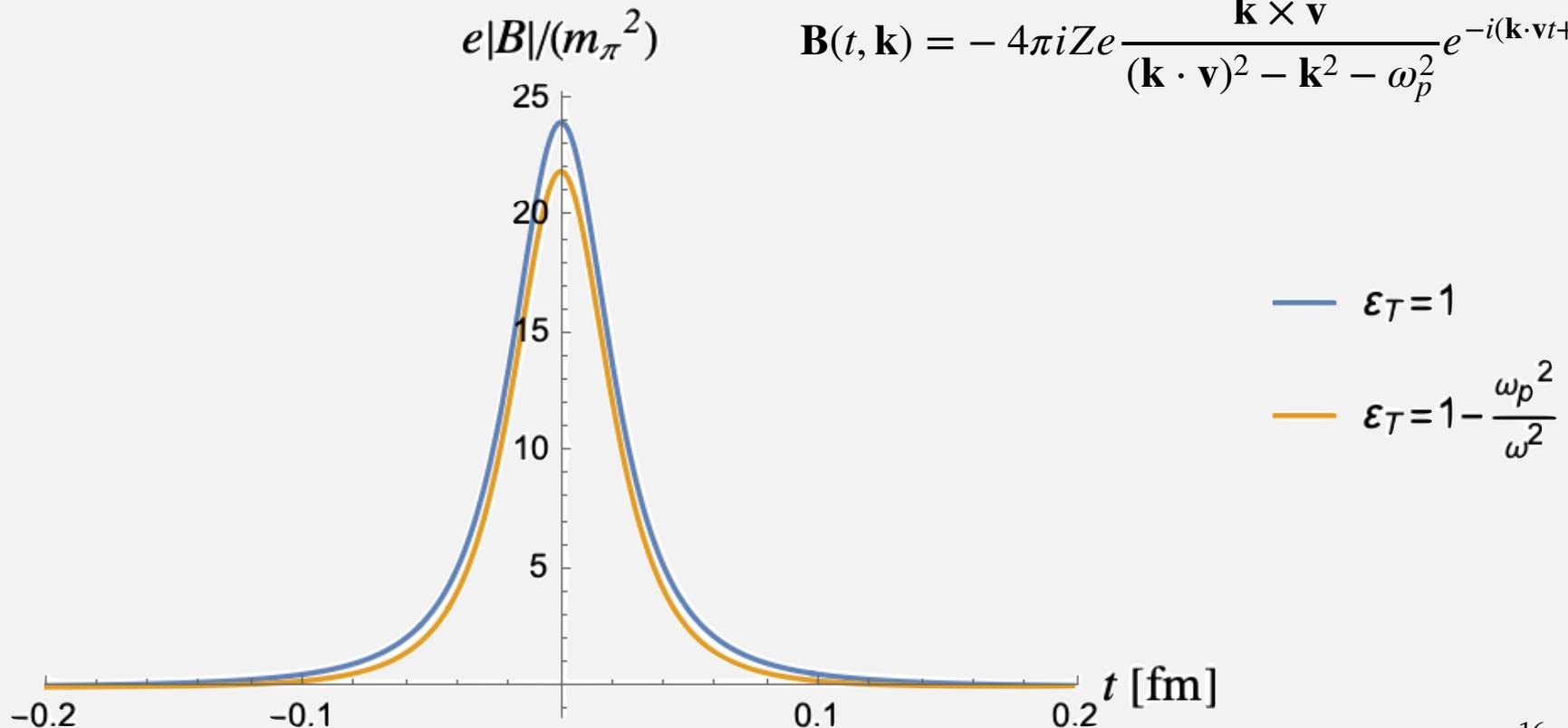
Magnetic field in the medium

- collisionless plasma -

$$\mathbf{B}(\omega, \mathbf{k}) = -4\pi i \frac{\mathbf{k} \times \mathbf{j}_{ext}(\omega, \mathbf{k})}{\omega^2 - \omega_p^2 - \mathbf{k}^2}$$

$$\varepsilon_T = 1 - \frac{\omega_p^2}{\omega^2}$$

$$\mathbf{B}(t, \mathbf{k}) = -4\pi i Z e \frac{\mathbf{k} \times \mathbf{v}}{(\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2 - \omega_p^2} e^{-i(\mathbf{k} \cdot \mathbf{v}t + \mathbf{k} \cdot \mathbf{r}_0)}$$



Magnetic field in the medium for the Au-Au collision

Magnetic field in the medium

- collisional plasma -

$$\varepsilon_T = 1 - \frac{\omega_p^2}{\omega(\omega + i\nu)}$$

$$\mathbf{B}(\omega, \mathbf{k}) = -4\pi i \frac{\mathbf{k} \times \mathbf{j}_{ext}(\omega, \mathbf{k})}{\omega^2 - \frac{\omega \omega_p^2}{\omega + i\nu} - \mathbf{k}^2}$$

$$\mathbf{B}(t, \mathbf{k}) = -4\pi i Z e e^{-i(\mathbf{k} \cdot \mathbf{v} t + \mathbf{k} \cdot \mathbf{r}_0)} \left[\frac{(\mathbf{k} \cdot \mathbf{v})^2 ((\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2 - \omega_p^2) + \nu^2 ((\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2)}{(\mathbf{k} \cdot \mathbf{v})^2 ((\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2 - \omega_p^2)^2 + \nu^2 ((\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2)^2} - i \frac{\nu (\mathbf{k} \cdot \mathbf{v}) \omega_p^2}{(\mathbf{k} \cdot \mathbf{v})^2 ((\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2 - \omega_p^2)^2 + \nu^2 ((\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2)^2} \right] \mathbf{k} \times \mathbf{v}$$

The magnetic field in heavy-ion collisions

The one-sided Fourier transformation

Magnetic field in reality

- Before the collision ($t < 0$) there is no plasma. Only the approaching nuclei generate the magnetic field.
- The additional fields generated due to the plasma appear at $t = 0$
- At $t > 0$ we see the effects of the quark-gluon plasma on the electromagnetic field.

$$f(\omega, \mathbf{k}) \equiv \int_0^{\infty} dt \int d^3r e^{i(\omega t - \mathbf{k} \cdot \mathbf{r})} f(t, \mathbf{r})$$

$$f(t, \mathbf{r}) = \int_{-\infty + i\sigma}^{\infty + i\sigma} \frac{d\omega}{2\pi} \int \frac{d^3k}{(2\pi)^3} e^{-i(\omega t - \mathbf{k} \cdot \mathbf{r})} f(\omega, \mathbf{k})$$

- The Maxwell's equations

$$\begin{cases} i\mathbf{k} \cdot \mathbf{D}(\omega, \mathbf{k}) = 4\pi\rho_{ext}(\omega, \mathbf{k}), \\ i\mathbf{k} \cdot \mathbf{B}(\omega, \mathbf{k}) = 0, \\ i\mathbf{k} \times \mathbf{E}(\omega, \mathbf{k}) = i\omega\mathbf{B}(\omega, \mathbf{k}) + \mathbf{B}_0(\mathbf{k}), \\ i\mathbf{k} \times \mathbf{B}(\omega, \mathbf{k}) = 4\pi\mathbf{j}_{ext}(\omega, \mathbf{k}) - i\omega\mathbf{D}(\omega, \mathbf{k}) - \mathbf{D}_0(\mathbf{k}). \end{cases}$$

The magnetic field and the initial fields

$$\mathbf{B}(\omega, \mathbf{k}) = -\frac{i}{\omega^2 \varepsilon_T(\omega, \mathbf{k}) - \mathbf{k}^2} \left[\mathbf{k} \times (4\pi \mathbf{j}_{ext}(\omega, \mathbf{k}) - \mathbf{E}_0(\mathbf{k})) - \frac{\mathbf{k}^2}{\omega} \mathbf{B}_0(\mathbf{k}) \right] + \frac{i}{\omega} \mathbf{B}_0(\mathbf{k})$$

$$\mathbf{j}_{ext}(\omega, \mathbf{k}) = i \frac{Ze \mathbf{v} e^{-i\mathbf{k} \cdot \mathbf{r}_0}}{\omega - \mathbf{k} \cdot \mathbf{v} + i0^+}$$

- The initial fields generated at $t = 0$

$$\mathbf{E}_0(\mathbf{k}) = -4\pi i \int \frac{d\omega}{2\pi} \left[\frac{\mathbf{k}(\mathbf{k} \cdot \mathbf{j}_{ext}(\omega, \mathbf{k}))}{\omega(\omega^2 - \mathbf{k}^2)} + \frac{\omega \mathbf{j}_{ext}(\omega, \mathbf{k})}{\omega^2 - \mathbf{k}^2} \right]$$

$$\mathbf{B}_0(\mathbf{k}) = -4\pi i \int \frac{d\omega}{2\pi} \frac{\mathbf{k} \times \mathbf{j}_{ext}(\omega, \mathbf{k})}{\omega^2 - \mathbf{k}^2}$$

$$\mathbf{j}_{ext}(\omega, \mathbf{k}) = 2\pi Ze \mathbf{v} \delta(\omega - \mathbf{k} \cdot \mathbf{v}) e^{-i\mathbf{k} \cdot \mathbf{r}_0}$$

- We use the **one-sided** Fourier transformation for the magnetic field, but the **two-sided** transformation for the initial fields! (The same for their current densities)

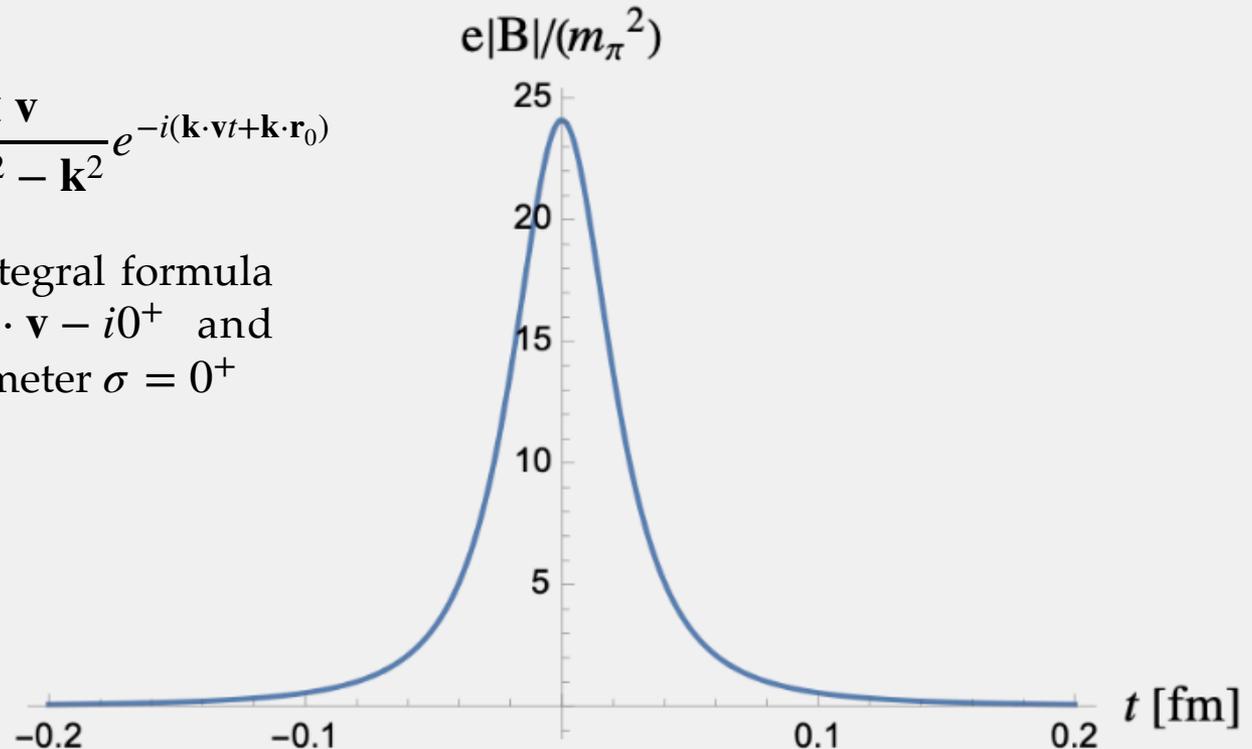
Magnetic field in the vacuum

$$\varepsilon_T = 1$$

$$\mathbf{B}(\omega, \mathbf{k}) = 4\pi Z e^{-i\mathbf{k}\cdot\mathbf{r}_0} \frac{\mathbf{k} \times \mathbf{v}}{(\omega^2 - \mathbf{k}^2)(\omega - \mathbf{k} \cdot \mathbf{v} + i0^+)} + i \frac{\mathbf{k} \times \mathbf{E}_0(\mathbf{k})}{\omega^2 - \mathbf{k}^2} + i \frac{\mathbf{k}^2 \mathbf{B}_0(\mathbf{k})}{\omega(\omega^2 - \mathbf{k}^2)} + \frac{i}{\omega} \mathbf{B}_0(\mathbf{k})$$

$$\mathbf{B}(t, \mathbf{k}) = -4\pi i Z e \frac{\mathbf{k} \times \mathbf{v}}{(\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2} e^{-i(\mathbf{k}\cdot\mathbf{v}t + \mathbf{k}\cdot\mathbf{r}_0)}$$

- We used the Cauchy's integral formula with poles at $\omega = \mathbf{k} \cdot \mathbf{v} - i0^+$ and $\omega = \pm |\mathbf{k}|$, and the parameter $\sigma = 0^+$



Magnetic field in the vacuum for the Au-Au collision

Magnetic field in the medium

$$\varepsilon_T = 1 - \frac{\omega_p^2}{\omega^2}$$

$$\mathbf{B}(\omega, \mathbf{k}) = 4\pi Z e^{-i\mathbf{k}\cdot\mathbf{r}_0} \frac{\mathbf{k} \times \mathbf{v}}{(\omega^2 - \mathbf{k}^2 - \omega_p^2)(\omega - \mathbf{k} \cdot \mathbf{v} + i0^+)} + i \frac{\mathbf{k} \times \mathbf{E}_0(\mathbf{k})}{\omega^2 - \mathbf{k}^2 - \omega_p^2} + i \frac{\mathbf{k}^2 \mathbf{B}_0(\mathbf{k})}{\omega(\omega^2 - \mathbf{k}^2 - \omega_p^2)} + \frac{i}{\omega} \mathbf{B}_0(\mathbf{k})$$

$$\begin{aligned} \mathbf{B}(t, \mathbf{k}) = & -i4\pi Z e^{-i\mathbf{k}\cdot\mathbf{r}_0} \left[\frac{e^{-i\sqrt{\mathbf{k}^2 + \omega_p^2} t}}{2\sqrt{\mathbf{k}^2 + \omega_p^2} (\sqrt{\mathbf{k}^2 + \omega_p^2} - \mathbf{k} \cdot \mathbf{v})} + \frac{e^{i\sqrt{\mathbf{k}^2 + \omega_p^2} t}}{2\sqrt{\mathbf{k}^2 + \omega_p^2} (\sqrt{\mathbf{k}^2 + \omega_p^2} + \mathbf{k} \cdot \mathbf{v})} \right. \\ & + \frac{e^{-i\mathbf{k}\cdot\mathbf{v}t}}{(\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2 - \omega_p^2} + \frac{e^{-i\sqrt{\mathbf{k}^2 + \omega_p^2} t} - e^{i\sqrt{\mathbf{k}^2 + \omega_p^2} t}}{2\sqrt{\mathbf{k}^2 + \omega_p^2}} \frac{\mathbf{k} \cdot \mathbf{v}}{(\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2} \\ & \left. + \left(-\frac{1}{\mathbf{k}^2 + \omega_p^2} + \frac{e^{-i\sqrt{\mathbf{k}^2 + \omega_p^2} t} + e^{i\sqrt{\mathbf{k}^2 + \omega_p^2} t}}{2(\mathbf{k}^2 + \omega_p^2)} \right) \frac{\mathbf{k}^2}{(\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2} + \frac{1}{(\mathbf{k} \cdot \mathbf{v})^2 - \mathbf{k}^2} \right] \mathbf{k} \times \mathbf{v} \end{aligned}$$

Future plans

- Finding a method to integrate highly oscillatory functions
- Calculating the magnetic field for the model more complex than that using point-like charges