

# Electromagnetic and Gravitational Hopfion-like solutions in de Sitter spacetime

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Joint work with J. Jezierski and A. Grzela

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# Motivation

- In Minkowski spacetime, Hopfions are finite energy, soliton-like family of solutions of Maxwell or linearized gravity theory

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- Closely related with Hopf fibration
- For quantum electrodynamics, pre-Hopfion solution treated like a photonic wave function saturates uncertainty relation for photons in three dimensions<sup>1</sup>:

$$\Delta r \Delta p \geq 4\hbar \quad (1)$$

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- Reduced data for pre-Hopfion is generated by imaginary shift in time for fundamental solution of wave equation:

$$\Phi_0 = (\mathbf{E} + i\mathbf{B}) \cdot \mathbf{r} = \frac{1}{(r^2 - (t - i)^2)} \quad (2)$$

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- Higher dimensional analog of Magic field—Imaginary shift in spatial direction  $z \rightarrow z - i$  for Coulomb solution  $1/R$  leads to

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- Example of simply but highly non-trivial solution of Maxwell equations — limit of Kerr–Newman E-M field<sup>2</sup> for mass  $\rightarrow 0$

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# Plan of the talk

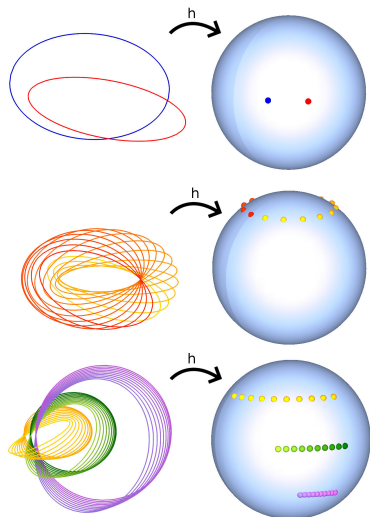
1. Brief introduction to Hopfions in Minkowski
  - Hopf fibration
  - Rañada's classical Hopfion
  
2. Hopfion-like solutions in de Sitter spacetime



# Hopf fibration

Hopf fibration is a non-trivial principal bundle of a three-dimensional sphere:

$$S^3 \sim \mathbb{R}^3 \cup \{\infty\} \xrightarrow{h} \mathbb{C} \cup \{\infty\} \sim S^2 \quad (3)$$

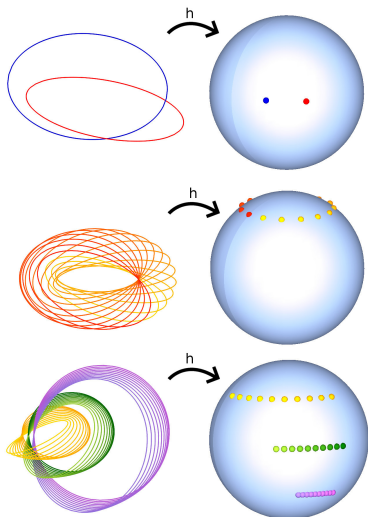


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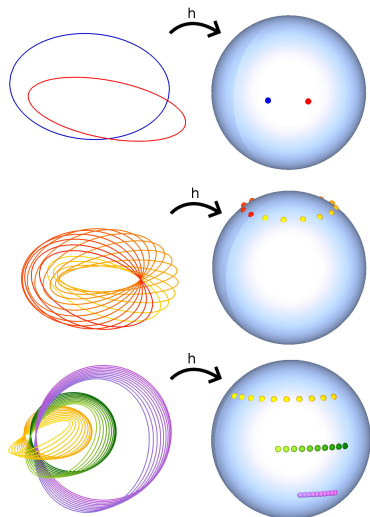
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The mapping  $h : \mathbb{R}^3 \cup \{\infty\} \xrightarrow{h} \mathbb{C} \cup \{\infty\}$  is given by:

$$h(x, y, z) = \frac{x + iz}{-y + i(\tilde{A} - 1)} \quad (4)$$

where  $\tilde{A} = \frac{1}{2}(x^2 + y^2 + z^2 + 1)$ .



# Rañada's classical Hopfion

In 1990, Rañada proposed the following solution of Maxwell equations:

$$\mathbf{E}_R = \frac{1}{4\pi} \frac{\nabla\xi \times \nabla\bar{\xi}}{(1 + \xi\bar{\xi})^2} \quad (5)$$

$$\mathbf{B}_R = \frac{1}{4\pi} \frac{\nabla\eta \times \nabla\bar{\eta}}{(1 + \eta\bar{\eta})^2} \quad (6)$$

where  $\xi(t, x, y, z)$  and  $\eta(t, x, y, z)$  are defined as

$$\xi = \frac{(Ax + ty) + \imath(Az + t(A - 1))}{(tx - Ay) + \imath(A(A - 1) - tz)} \quad (7)$$

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Properties:

■  $\xi(0, x, y, z) = h(x, y, z)$  and  $\eta(0, x, y, z) = h(z, x, -y)$

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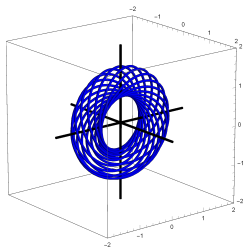
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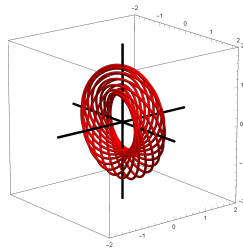
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- The helicities are preserved in time.

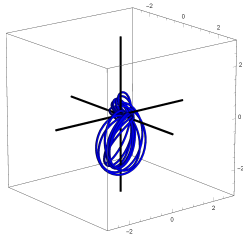
# Field lines



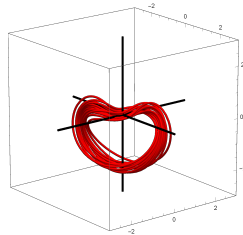
Hopfions ( $L=1$ ), electric,  $t=0$



Hopfions ( $L=1$ ), magnetic,  $t=0$



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Hopfions ( $L=1$ ), magnetic,  $t=1$



# Topological invariants in physics

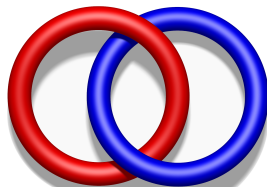
A topologically invariant configuration of field lines is to be linked and/or knotted.

A measure of linkedness is Gauss linking integral:

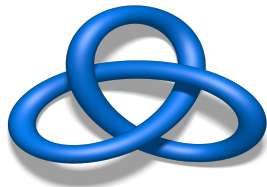
$$L(\mathbf{c}_1, \mathbf{c}_2) = \frac{1}{4\pi} \int \frac{d\mathbf{c}_1}{dt_1} \cdot \frac{\mathbf{c}_1 - \mathbf{c}_2}{\|\mathbf{c}_1 - \mathbf{c}_2\|^3} \times \frac{d\mathbf{c}_2}{dt_2} \quad (9)$$

where  $\mathbf{c}_1(t_1)$ ,  $\mathbf{c}_2(t_2)$  are closed curves.

The self-linking number,  $L(\mathbf{c}, \mathbf{c})$ , is a measure of knottedness.



Linked,  $L=1$ .



Knotted

# Topological invariant in physics

Physical analogue of Linking number are helicities<sup>3</sup>. The magnetic helicity:

$$h_m = \int_V dV \mathbf{A} \cdot \mathbf{B} \quad (10)$$

where  $\mathbf{B} = \text{rot } \mathbf{A}$ .

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Analogously we can define electric helicity

$$h_e = \int_V dV \mathbf{C} \cdot \mathbf{E} \quad (11)$$

where  $\mathbf{E} = \text{rot } \mathbf{C}$ .

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# Helicity as a Noether current

## ■ Lagrangian for Maxwell theory

$$\mathcal{L}_{\text{EM}} = -\frac{1}{4}F^{\alpha\beta}F_{\alpha\beta} = \frac{1}{2}(E^2 - B^2) \quad (12)$$

where  $E^2 = E_i E^i$ ,  $B^2 = B_i B^i$ .

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$$\partial_\beta F^{\alpha\beta} = 0 \quad \partial_\beta (*F)^{\alpha\beta} = 0 \quad (13)$$

is invariant under the duality reflection  $F_{\alpha\beta} \rightarrow (*F)_{\alpha\beta}$ .

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## ■ Following<sup>4</sup>, we analyze the duality symmetric Lagrangian

$$\mathcal{L}_{\text{EM-ds}} = -\frac{1}{8} \left( F_{\alpha\beta} F^{\alpha\beta} + G_{\alpha\beta} G^{\alpha\beta} \right) \quad (14)$$

where we impose duality constraint  $G_{\alpha\beta} = (*F)_{\alpha\beta}$ .

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$$F_{\alpha\beta} \rightarrow F_{\alpha\beta} \cos \theta + G_{\alpha\beta} \sin \theta \quad G_{\alpha\beta} \rightarrow G_{\alpha\beta} \cos \theta - F_{\alpha\beta} \sin \theta \quad (15)$$

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- Corresponding Noether current, known as helicity current

$$J_{\mathcal{H}}^{\alpha} = \frac{\partial \mathcal{L}_{\text{EM-ds}}}{\partial (\partial_{\alpha} A_{\beta})} C_{\beta} - \frac{\partial \mathcal{L}_{\text{EM-ds}}}{\partial (\partial_{\alpha} C_{\beta})} A_{\beta} = \frac{1}{2} (G^{\alpha\beta} A_{\beta} - F^{\alpha\beta} C_{\beta}) \quad (16)$$

where the helicity density  $\mathcal{H}$  and helicity flux density—the spin density  $\mathbf{J}_{\mathcal{H}}$  are, in transverse gauge

$$J_{\mathcal{H}}^0 \equiv \mathcal{H} = \frac{1}{2} (\mathbf{A} \cdot \mathbf{B} - \mathbf{C} \cdot \mathbf{E}) \quad \mathbf{J}_{\mathcal{H}} = \frac{1}{2} (\mathbf{E} \times \mathbf{A} + \mathbf{B} \times \mathbf{C}) \quad (17)$$

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- Provided the equation of motion holds, the helicity current is conserved  $\partial_{\alpha} J_{\mathcal{H}}^{\alpha} = 0$ ,

$$\dot{\mathcal{H}} + \nabla \cdot \mathbf{S} = 0 \quad (18)$$

# Conformal flatness of de Sitter spacetime

De Sitter spacetime equipped with the metric

$$g = -(1 - k^2 r^2) dt^2 + \frac{dr^2}{1 - k^2 r^2} + r^2 (d\theta^2 + \sin^2 \theta d\phi^2) \quad (19)$$

can be locally transformed in a conformally flat metric in Cartesian coordinates

$$g_{dS} = \frac{4}{k^2 (1 + S^2)^2} (-dT^2 + dX^2 + dY^2 + dZ^2) \quad (20)$$

where  $S^2 = -T^2 + X^2 + Y^2 + Z^2$ . Exact form of the transform reads

$$\begin{aligned} T &= \frac{\sqrt{1-k^2 r^2} \sinh kt}{1 - \sqrt{1-k^2 r^2} \cosh kt}, & X &= \frac{kr \sin \theta \cos \phi}{1 - \sqrt{1-k^2 r^2} \cosh kt}, \\ Y &= \frac{kr \sin \theta \sin \phi}{1 - \sqrt{1-k^2 r^2} \cosh kt}, & Z &= \frac{kr \cos \theta}{1 - \sqrt{1-k^2 r^2} \cosh kt} \end{aligned} \quad (21)$$

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$$Z^r = \frac{4e^{i\phi} \sin \theta s(r)}{\left(\left(1 - \frac{4}{k^2}\right) s(r) + \frac{4}{k^2} + 1\right)^2} \quad (22)$$

$$Z^\theta = \frac{4k^4 e^{i\phi} \left(\cos \theta \left(k^2 (s(r) + 1) + 4(s(r) - 1)\right) - 4ik^2 r s^3(r)\right)}{r \left(k^2 (s(r) + 1) - 4s(r) + 4\right)^3} \quad (23)$$

$$Z^\phi = -\frac{4k^4 e^{i\phi} \left(-4k^2 r \cos \theta s^3(r) - i \left(k^2 (s(r) + 1) + 4(s(r) - 1)\right)\right)}{r \sin \theta \left(k^2 (s(r) + 1) - 4s(r) + 4\right)^3} \quad (24)$$

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- Weak  $k$  limit

$$\mathbf{Z} = \mathbf{Z}_{\text{Minkowski}} + O(k^2) \quad (25)$$

## ■ Near cosmological horizon analysis

$$\begin{aligned}
 Z^r &= \frac{4\sqrt{2k}e^{i\phi} \sin \theta}{\left(1 + \frac{4}{k^2}\right)^2} \sqrt{\frac{1}{k} - r} + O\left(\frac{1}{k} - r\right) \\
 Z^\theta &= \frac{4k^5 (k^2 - 4) e^{i\phi} \cos \theta}{(k^2 + 4)^3} + O\left(\sqrt{\frac{1}{k} - r}\right) \\
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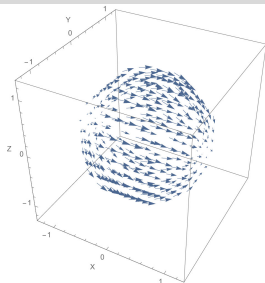
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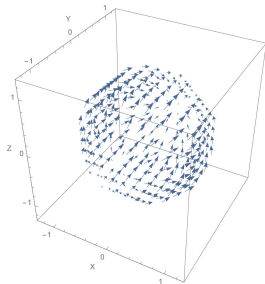
$$Z^\theta = \frac{4k^5(k^2 - 4)e^{i\phi}\cos\theta}{(k^2 + 4)^3} + O\left(\sqrt{\frac{1}{k} - r}\right)$$

$$Z^\phi = \frac{4ik^5(k^2 - 4)e^{i\phi}}{\sin\theta(k^2 + 4)^3} + O\left(\sqrt{\frac{1}{k} - r}\right)$$

- The fields flow from one pole of the sphere to the other along the OX (electric) and OY (magnetic) axes.



**Electric field**



**Magnetic field**

# Field lines

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- For all  $x$ :  $E_y(x, 0, 0) = E_z(x, 0, 0) = 0$
- We obtain a straight section with a parametrization

$$r'(\tau) = \frac{4\sqrt{1 - k^2 r(\tau)^2}}{\left( \left(1 - \frac{4}{k^2}\right) \sqrt{1 - k^2 r(\tau)^2} + \frac{4}{k^2} + 1 \right)^2} \quad (26)$$

## Field lines

- Consider a geodesic for electric field with initial condition  $(x, y, z) = (x_0 > 0, 0, 0)$
- For all  $x$ :  $E_y(x, 0, 0) = E_z(x, 0, 0) = 0$
- We obtain a straight section with a parametrization

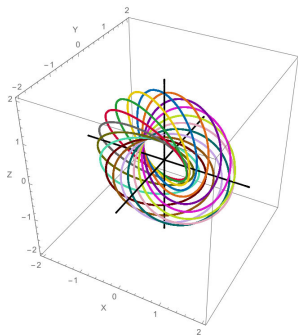
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- It reach cosmological horizon  $r = 1/k$  in finite time

$$\tau_H = \frac{(8 + 3\pi)k^4 + 8\pi k^2 + 16(3\pi - 8)}{16k^5} \quad (27)$$

# Integral curves of $B$ field, $t = 0$ — numeric analysis

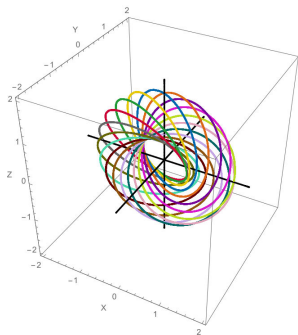
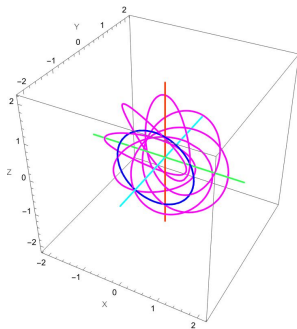
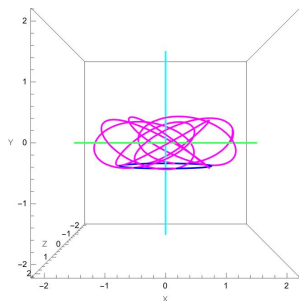
- Part of Integral curves behaves like in the Minkowski case (Hopf fibration structure) – fig. (a)



(a)  $k = 10^{-6}$

# Integral curves of $B$ field, $t = 0$ — numeric analysis

- Part of Integral curves behaves like in the Minkowski case (Hopf fibration structure) – fig. (a)
- Group of closed field lines which cross XZ plane many times – fig. (b)

(a)  $k = 10^{-6}$ (b.1)  $k = 0.32$ (b.2)  $k = 0.32$

# Summary

- Hopfions are real (finite energy) soliton-like family of solutions of Maxwell (linearized gravity) theory

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<sup>5</sup>T. S. J. Jezierski *Simple description of generalized electromagnetic and gravitational hopfions*, (2018) *Class. Quantum Grav.* **35** 245010.



# Summary

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- We propose Hopfion-like solutions for de Sitter spacetime which become Hopfions in  $k \rightarrow 0$  limit

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  - Field lines which reach cosmological horizon in finite time parameter
  - Closed field lines for which linking number is greater than one. Various cases in point of view of topological classes
- In our paper<sup>5</sup>, presented construction enables one to obtain corresponding solution for Linearized Gravity

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- 1 W. T. Irvine, D. Bouwmeester *Linked and knotted beams of light* Nature Physics, **4** 716-720, (2008).
- 2 M. Arrayás, D. Bouwmeester, J. L. Trueba *Knots in electromagnetism* Physics Reports 667 (2017)
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Thank You!