

NATIONAL CENTRE FOR NUCLEAR RESEARCH ŚWIERK

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INCONTRECT

Searches for Gravilational Waves from Known Pulsars at Two Harmonics in the second and third LIGO-Virgo Observing Runs (Abbott et al 2022 ApJ 935 1)





PULSATS

- Pulsars are a special kind of neutron stars. - They emit a beam of radiation which we observe as a periodic pulse as the beam sweeps over us once per rotation.

Particles accelerated along strong magnetic field lines produce beam of radiation. If magnetic field is not aligned the spin axis, this gives lighthouse effect.



Courtesy: Manchester, R.N and Taylor, J.H., Pulsars, Freeman, 1977.



- 1934: Walter Baade and Fritz Zwicky predicted the existence of NS

- 1967: First pulsar was discovered by Jocelyn Bell and Anthony Hewish.





Courtesy: arxiv 1508.03115

- Young and strongly magnetized pulsars: Top right hand corner

- Old and weakly magnetized pulsars: Bottom left

Estimated magnetic field: $B = 3.2 \times 10^{19} \sqrt{P\dot{P}} G$

Standard pulsar: $B \approx 10^{12} G$

Accreting pulsar: $B \approx 10^8 G$

Magnetar: $B \approx 10^{15} G$

A rough estimate of pulsar's age: $\tau_c = \frac{P}{2\dot{P}}$

The angular velocity in the Z'-direction produces angular momentum in all three directions. In this case; $f_{\rm GW} = f_{\rm rot} \ (l = 2, m = 1)$ and

 $h_{21} = -\frac{C_{21}}{2} \left[F_+^D(\alpha, \delta, \psi; t) \sin \iota \cos \iota \cos \left(\Phi(t) + \Phi_{21}^C \right) + F_\times^D(\alpha, \delta, \psi; t) \sin \iota \sin \left(\Phi(t) + \Phi_{21}^C \right) \right]$ $h_{22} = -C_{22} \left[F_{+}^{D}(\alpha, \delta, \psi; t)(1 + \cos^{2} \iota) \cos\left(2\Phi(t) + \Phi_{22}^{C}\right) + 2F_{\times}^{D}(\alpha, \delta, \psi; t) \cos\iota \sin\left(2\Phi(t) + \Phi_{22}^{C}\right) \right]$ C_{21} and C_{22} : dimensionless constants α and δ : right ascension and declination of the source i and ψ : describe the orientation of the source's spin axis with respect to the observer in terms of inclination and polarization and Φ_{22}^C : phase angles at a defined epoch Φ_{21}^{C} $\Phi(t)$: rotational phase of the source

Dual Harmonics

 $L_x = 2\pi f_{\rm rot} I_{xz}$ $L_{y} = 2\pi f_{\rm rot} I_{yz}$ $L_{7} = 2\pi f_{\rm rot} I_{77}$

1
$$f_{\rm GW} = 2f_{\rm rot} \ (l = 2, m = 2)$$

 $\vec{L} = \begin{bmatrix} I_{xx} & 0 & 0 \\ 0 & I_{yy} & 0 \\ 0 & 0 & I_{77} \end{bmatrix} \begin{bmatrix} 0 \\ 0 \\ 2\pi f_{rot} \end{bmatrix}$

direction. In this case; $f_{\text{GW}} = 2f_{\text{rot}}$ (l = 2, m = 2)

 $C_{21} = 0$ and $C_{22} = 2h_0$ where $h_0 = \frac{16\pi^2 G I_{zz} \varepsilon f_{rot}^2}{c^4}$,

 ε is the ellipticity defined as $\varepsilon \equiv \frac{|I_{xx} - I_{yy}|}{I_{xx}}$ d is the distance of the source $\left(\frac{d}{dt}\left(\frac{1}{2}I_{zz}\Omega^{2}\right) = \frac{G}{5c^{5}} < \ddot{Q}_{m}^{ij}(t')\ddot{Q}_{m}^{ij}(t') > \right)$ gives

 $L_x = L_y = 0$

 $L_7 = 2\pi f_{\rm rot} I_{77}$

The angular velocity in the Z'-direction produces angular momentum only in the Z'-

 ${\cal E}$

 $f_{GW} = 2f_{rot}$

spin-down limit: $h_0^{\rm sd} = \frac{1}{d} \left(\frac{5GI_{zz}}{2c^3} \frac{|\dot{f}_{\rm rot}|}{f_{\rm rot}} \right)$

Riemann Cheorem:

No. of DOF in N dimensional metric = N(N-1)

FIG. 1. Effect of different GW polarizations on a ring of free-falling test particles. Plus (+) and cross (\times) tensor modes (green); vector-x (x) and vector-y (y) modes (red); breathing (b) and longitudinal (l) scalar modes (black). In all of these diagrams the wave propagates in the z direction. This decomposition into polarizations was first proposed for generic metric theories in [7].

Courtesy: LIG0-P1600305

gives

Tensor modes, s = 2Vector modes, s = 1Scalar modes, s = 0

4 dimensional metric

Why do we search for non-GR polarizations?

- Hubble tension - QFT in curved space time

"...there are the ones that invent OCCULT" FLUIDS to understand the Laws of Nature. They will come to conclusions, but they now run out into DREAMS and CHIMERAS neglecting the true constitution of things..... ...however there are those that from the simplest observation of Nature, they reproduce New Forces (i.e. New Theories)... "

From the Preface of PRINCIPIA (II Edítíon) 1687 by Isaac Newton, wrítten by Mr. Roger Cotes

Courtesy: S. Capozziello

BD theory: 2 tensor states (similar to GR) and 1 scalar state We assume a spherical star with a tiny mountain on its equator. This mountain can be approximated by a Dirac delta function. In this case: $f_{\rm GW} = f_{\rm rot} \ (l=1, m=1)$

The signal h_{11} due to the dipole radiation is given by:

 $c(\alpha, \delta; t)$: amplitude modulation function 5 is the BD parameter obtained observationally h_0^d : Signal amplitude given by $h_0^d = \frac{4\pi G}{c^3} \zeta \frac{Df_{\rm rot}}{d}$

The spin-down limit for the dipole radiation is

$$\frac{d}{dt}\left(\frac{1}{2}I\Omega^{2}\right) = \frac{c^{3}}{16\pi G\zeta}\omega^{2}r^{2}\left(h_{0}^{d}\right)^{2} < \cos^{2}(\omega t) > \int_{\iota=0}^{\pi}\int_{\rho=0}^{2\pi}\sin^{3}\iota d\iota d\rho$$

Dipole radiation in Brans-Dicke (BD)

- $h_{11} = -h_0^d c(\alpha, \delta; t) \sin \iota \sin(\Phi(t) + \Phi_0)$

Plus polarization (h_+)

 $h_{+} \& h_{\times}$:

 h_S :

- Tensor polarizations in GR

- Dominated by time-varying quadrupole moment
- 2-fold symmetric
- Ellipses preserve the area

Scalar polarization in BD - Dominated by time-varying dipole moment

- co-fold symmetric

Colariza Elons

Cross polarization (h_x)

Scalar polarization (h_S)

Electromagnetic data using CHIME, Hobart, Jodrell Bank, MeerKAT, Nancay, NICER and UTMOST observatories.

Oct 30, 2016

April 01, 2019

During 03, the Virgo detector was operational for the entire period.

03

Aug 25, 2017

Virgo (Aug 01, 2017)

March 27, 2020

1. Time domain Bayesian Method

- Fermi Dirac distribution priors on amplitude parameters are used
- Raw GW data is heterodyned using their expected phase evolution
- relativistic effects

2. Sn-Vector Method

- Multi-detector matched filter in the frequency domain - Based on Sidereal modulation of the expected signal amplitude and phase

Search Melhods

- Includes correction for relative motion of source wrt detectors and various

3. F/G/D-statistics

F-statistic:

G-statistic: amplitude & phase

D-scalislic:

Dipole radiation in BD theory

amplitude, phase and polarizations

polarizations

known

Bayesian analysis	Sn-Vector Method	\$
Single Harmonic		singl
Dual Harmonic	L=2, m=1	Dua
		Dipol

Only for 23 high value pulsars

Analysis

$\mathcal{F}/\mathcal{G}/\mathcal{D}$ latistics

le harmonic

Harmonic

le radiation

- Total no. of pulsars = 236

- Total no. of pulsars in binaries = 168
- No. of pulsars with $f_{rot} > 100 \text{ Hz}$ = 161
- No. of high value pulsars $(h_0 < h_0^{sd}) = 23$

Abbolt et al 2022 ApJ 935 1

RESULES

- Grey triangle: spin-down limit for each pulsar
- Shaded circles: ho for pulsars which surpass the spin-down limits.
- Green dotted line: If upper limit < spindown Limit.
- Solid line: Joint detector sensitivity estimate for 03

Abbott et al 2022 ApJ 935 1

Crab (J0534+2200 Glitch occurred GWs contributes down $h_0^{95\%} = 1.3(1.2) \times$ $Q_{22}^{95\%} = 5.6(5.0) \times \epsilon^{95\%} = 7.2(6.5) \times \epsilon^{95\%}$

orientation.

Results (conta)

Limit on mass quadrupole moment which in turn imposes limit on ellipticity.

10 ⁻⁶	$\begin{pmatrix} h_0 \end{pmatrix}$	$\begin{pmatrix} d \end{pmatrix}$	$\left(100 Hz \right)^2$	$\left(\begin{array}{c} 10^{38} \ kg \ m^2 \end{array}\right)$
	(10^{-25})	$\left(1 \ kpc\right)$	f _{rot}	I_{zz}^{fid}

0)	Vela (J0835-4510)
	No glitch occurred
s < 0.009% of the spin-	GWs contributes $< 0.27\%$ of the spin-
	down
< 10 ⁻²⁶ at d = 2 kpc	$h_0^{95\%} = 1.8(1.7) \times 10^{-25}$ at d = 0.28 kpc
$ imes 10^{32} \text{kg-m}^2$	$Q_{22}^{95\%} = 7.2(7.1) \times 10^{33} \text{ kg-m}^2$
10 ⁻⁶	$\epsilon^{95\%} = 9.3(9.2) imes 10^{-5}$

 $h_0^{95\%}$ is the 95% credible upper limits on the amplitudes h_0 , $Q_{22}^{95\%}$ is the 95% credible upper limits on the quadrupole moments and $e^{95\%}$ is the 95% credible upper limits on the ellipticity. Values in parentheses are those produced using the restricted

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Dziekuje Bardzo.

